Mathematical Clarification of General Relativistic Variational Principles

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Abstract

In this paper a mathematically precise global (i.e. not the usual local) approach is presented to the variational principles of general relativistic classical field theories.

Problems of the classic (usual) approaches are also discussed in comparison.

The aim of developing a global approach is to provide a possible tool for future efforts on proving global existence theorems of field theoretical solutions.

1 Introduction

As one can find out from physics and mathematics literature, the known variational formulations of general relativistic classical field theories can be divided into three classes. These classes differ in the definition of the action functional (in the definition of the integration domain of the action functional), and in the notion of variation.¹

The first class of approaches specify the action functional as an integral of a Lagrange form² over the spacetime manifold. This approach is mathematically ill defined, as the action would diverge for some quite physical (e.g. stationary) field configurations, in general. (This fact can be shown explicitly on specific examples.)

The second class specifies the action as the integral of the Lagrange form on a given compact subset of the spacetime manifold. The variation is then defined by using one-parameter families of field configurations (see e.g. [5], [8]). This means a Gateaux-like notion of derivative.

The third class defines the action functional as the integral of the Lagrange form on the spacetime domain between two given time-slice. As time-slices may be noncompact in general, certain fall-off properties have to be introduced on the field configurations in order to make sense of the action (otherwise, the action could diverge for specific configurations). After specifying appropriate regularity conditions, one can define a natural C^k -type supremum norm equivalence class (for some nonnegative integer k), and the variation is simply defined as the (Fréchet) derivative of the action functional (with respect to this norm equivalence class).

¹The concept of variation of the action functional is a notion of a kind of derivative. Some of the approaches use one-parameter families of field configurations to define the variation of the action functional. This derivativelike notion resembles to the Gateaux derivative (directional derivative). Other approaches use more adequate notion for the variation, which corresponds to the Fréchet derivative (classic derivative of a map between a normed affine space and a topological vector space, based on the notion of ordo functions).

²We call the product of the Lagrangian density and the volume form the Lagrange form.

All the three approaches have certain undesired properties. The first formulation, as pointed out, is mathematically ill defined. The second one is a local kind of formulation, and the notion of variation is Gateaux-like. The Gateaux-like derivative is a much weaker notion than the Fréchet derivative: the most powerful tools in differential calculus (e.g. the Taylor formulas, some critical point theorems etc.) only work with Fréchet type derivative.³ Apart from this argument of mathematical inelegance, we will discuss a further problem of this approach in *Section 3.2.2* (the argument of non-constructiveness), which is in connection with boundary value problems. The third approach is semi-global, as it is global in spatial sense, but local in the timelike direction, furthermore the notion of variation is Fréchet type, which is potentially more powerful, when trying to prove e.g. some critical point theorems. However, there are several difficulties concerning this kind of formulation, which are discussed in the end of *Subsection* 3.2.2 (the problem of spacetime splitting and the problem of spatial fall-off conditions).

Inspired by the above mentioned problems, we developed a global kind of approach, which uses Fréchet type notion of variation. The domain of integration can be viewed as the conformal compactification of the arising spacetime model, and the notion of variation is simply the (Fréchet) derivative of the action functional, with respect to the natural C^k -type supremum norm equivalence class (for some nonnegative integer k) on the field configurations.

The presented approach resembles to Palatini formulation of general relativity, as the covariant derivation is taken to be an independent dynamical field quantity. The formalism can also handle theories using covariant derivations with nonvanishing torsion. A main result concerning theories with nonvanishing torsion can be found in *Section 4*, in *Theorem 15* (which presents the field equations and boundary constraints for the case of nonzero torsion).

The class of spacetime manifolds, which can be generated by this formulation, are the orientable, asymptotically simple models⁴.

A global approach, which uses Fréchet type notion of variation, can possibly be applied in future as a tool of proving global existence theorems in general relativistic field theories.

2 Building up a field theory by using variational principles

2.1 Base manifold

Let us take a real C^3 manifold M, which is orientable. M will be called the *base manifold*.⁵ Let us use the following notation: $m := \dim M$. In the followings, we will denote by \mathbb{R} the real numbers, and by \mathbb{N} the positive integers.

³It is also a well known fact, that the Fréchet differentiability property of a function at a point is a much stronger condition than the Gateaux (directional) differentiability. Although, in finite dimensions, the continuous Gateaux differentiability on an open set is known to be equivalent to the continuous Fréchet differentiability, this is not true in infinite dimensions (i.e. in our case).

⁴In the definition of asymptotic simpleness, we do not include the condition of asymptotic emptiness.

⁵The base manifold is going to be the manifold, where the integration is carried out in order to define the action functional. In the classic formalism, the base manifold plays the role of spacetime manifold. However, in *Section* 4 and 5, we shall carry out an argument that the base manifold should not be directly interpreted as the spacetime manifold, but as the so called conformal compactification of the arising spacetime model. Thus, we do not refer to the base manifold as spacetime manifold.

2.2 Field quantities

As usually in a classical field theory, the field quantities will be sections of a fixed real vector bundle over the base manifold, and covariant derivations over the vector bundle.⁶

If W(N) is a real $C^{\check{k}}$ ($\check{k} \in \{0, \ldots, k\}$) vector bundle over a real C^k manifold N, then the C^l ($l \in \{0, \ldots, \check{k}\}$) sections of it will be denoted by $\Gamma^l(W(N))$. Furthermore let $D^l(W(N))$ be the space of C^l covariant derivations over W(N). Let X(N) be a real $C^{\check{k}}$ ($\check{k} \in \{0, \ldots, k\}$) vector bundle over N. We use the natural injection of $D^{\check{l}}(W(N)) \times D^{\bar{l}}(X(N))$ into $D^l(W(N) \times X(N))$ ($\check{l}, \bar{l} \in \{0, \ldots, \min(\check{k}, \bar{k})\}$ and $l \in \{0, \ldots, \min(\check{l}, \bar{l})\}$).

Remark 1. Let F(N) be $N \times \mathbb{R}$ as a real C^k vector bundle. Let T(N) be the tangent vector bundle of N as a real C^k vector bundle. Let W(N) be a real $C^{\check{k}}$ ($\check{k} \in \{0, \ldots, k\}$) vector bundle. A C^l ($l \in \{0, \ldots, \check{k}\}$) covariant derivation in $D^l(T(N)) \times D^l(W(N))$ can be uniquely extended to a C^l covariant derivation over all the mixed tensor and cross products of F(N), T(N), W(N) and their duals by requiring the Leibniz rule, the commutativity with contraction, and the correspondence to the exterior derivation on F(N). We can refer to this unique extension by the original covariant derivation, because they determine each other uniquely. For example if $s \in \Gamma^2(W(N))$ and $\nabla \in D^2(T(M)) \times D^2(W(N))$, then $\nabla(\nabla(s))$ is well defined, if k and \check{k} is greater or equal to 2.

The C^l covariant derivations in $D^l(T(N)) \times D^l(W(N))$ form an affine space over the vector space of $\Gamma^{l-1}(T^*(N) \otimes ((T(N) \otimes T^*(N)) \times (W(N) \otimes W^*(N))))$ sections, that is over the vector space of the so called C^{l-1} diagonal Christoffel tensor fields of $T(N) \times W(N)$. This means, that if we subtract two such covariant derivation, their action corresponds to the action of a C^{l-1} Christoffel tensor field of T(N) (that is, to a C^{l-1} section of $T^*(N) \otimes T(N) \otimes T^*(N)$) on the sections of T(N), and to the action of a C^{l-1} Christoffel tensor field of W(N) (that is, to a C^{l-1} section of $T^*(N) \otimes W(N) \otimes W^*(N)$) on the sections of W(N). This fact follows from the basic properties of the covariant derivations: we refer to textbooks e.g. [5] and [8].

Let us fix a C^3 real vector bundle V(M) over M. Let us introduce a sub fiber bundle $\check{V}(M)$ of the vector bundle V(M), with the same fiber dimension as V(M) (thus, for each $p \in M$ the fiber $\check{V}_p(M)$ is a sub manifold of $V_p(M)$ with dimension $\dim(V_p(M))$). Let $\check{D}^3(T(M), V(M))$ be a closed sub affine space of the affine space $D^3(T(M)) \times D^3(V(M))$, where the topology is understood to be the topology defined in *Definition 8* in *Subsection 3.1*. The *field variables* of the theory are going to be the elements of $\Gamma^3(\check{V}(M)) \times \check{D}^3(T(M), V(M))$, that is the covariant derivations are also dynamical.

2.3 The Lagrange form

We introduce a central notion of the variational principles: the Lagrange form. It is going to replace the notion of Lagrangian density function of the classic formalism. This notion is well known in literature, but there is no generally accepted label for it. (In the classic formalism, the Lagrange form can be obtained as the product of the Lagrangian density and the volume form.)

 $^{^{6}}$ This kind of notion of field quantities is used in the Palatini type formulation of general relativistic variational problems. In our approach, the torsion of the covariant derivation is not assumed to be zero *a priori*.

Let us take a map

$$\begin{aligned} \mathbf{dL}: \Gamma^3\Big(\breve{V}(M)\Big) \times \Gamma^2\Big(T^*(M) \otimes V(M)\Big) \times \Gamma^2\Big(\stackrel{2}{\wedge} T^*(M) \otimes \big((T^*(M) \otimes T(M)) \times (V^*(M) \otimes V(M))\big)\Big) \\ & \to \Gamma^2\Big(\stackrel{m}{\wedge} T^*(M)\Big) \end{aligned}$$

which is *pointwise*, that is

Given such a map \mathbf{dL} , for every $p \in M$ we can naturally define the map

$$\mathbf{dL}_p: \breve{V}_p(M) \times \left(T_p^*(M) \otimes V_p(M)\right) \times \left(\bigwedge^2 T_p^*(M) \otimes \left((T_p^*(M) \otimes T_p(M)) \times (V_p^*(M) \otimes V_p(M))\right)\right) \\ \to \bigwedge^m T_p^*(M)$$

with the restriction of \mathbf{dL} (this new function maps between finite dimensional vector spaces). If for every $p \in M \mathbf{dL}_p$ is C^2 , then we call \mathbf{dL} a Lagrange form. (The above requirements mean, that a Lagrange form can also be viewed as a C^2 fiber bundle homomorphism.)

Remark 2. Let us take a Lagrange form **dL**. Let us denote the partial derivative of \mathbf{dL}_p in its r-th variable $(r \in \{1, 2, 3\})$ by $D_r \mathbf{dL}_p$ $(p \in M)$. Let us take any section

$$(v,w,r) \in \Gamma^{3}(\breve{V}(M)) \times \Gamma^{2}(T^{*}(M) \otimes V(M)) \times \Gamma^{2} \Big(\bigwedge^{2} T^{*}(M) \otimes \big((T^{*}(M) \otimes T(M)) \times (V^{*}(M) \otimes V(M)) \big) \Big).$$

Then the derivative $D_r \mathbf{dL}_p(v_p, w_p, r_p)$ $(r \in \{1, 2, 3\})$ can be viewed as a tensor at $p \in M$ of the appropriate type, because it is a linear map between the appropriate vector bundle fibers at $p \in M$. Furthermore, the tensor field defined by $p \mapsto D_r \mathbf{dL}_p(v_p, w_p, r_p)$ $(r \in \{1, 2, 3\})$ is C^1 . This follows from the following facts. The map \mathbf{dL}_p is C^2 for every $p \in M$. Furthermore, the map $p \mapsto \mathbf{dL}_p(v_p, w_p, r_p)$ is C^2 for every (v, w, r) section as above. Therefore, by taking a coordinate chart on an open subset of M and trivializations of the appropriate vector bundles over it, we see that the function \mathbf{dL} (taken in coordinates) is C^2 in its manifold coordinate variable, and is C^2 in its vector bundle fiber coordinate variable. Thus, we have that \mathbf{dL} (taken in coordinates) possesses the C^2 property.⁷ Therefore, any of its partial derivatives are C^1 : for example the partial derivative, which corresponds to $D_r \mathbf{dL}(r \in \{1, 2, 3\})$ is also C^1 (in coordinates). By this fact, the C^1 property of the tensor field $p \mapsto D_r \mathbf{dL}_p(v_p, w_p, r_p)$ ($r \in \{1, 2, 3\}$) is implied over an arbitrary coordinate neighborhood, thus on the whole manifold.

2.4 The action functional

The central notion of variational principles is the action functional. If one is trying to find an elegant formulation of classical field theories, the key step is the proper definition of the

⁷This follows from the following theorem: given a map between a finite product of finite dimensional vector spaces and a finite dimensional vector space, then it is C^1 if and only if it is partially C^1 all in its variables.

action functional. The action is the integral of the Lagrange form on the base manifold or on a properly specified subset of it.⁸

As a Lagrange form dL is volume form field valued, given a section

$$(v,w,r) \in \Gamma^{3}(\breve{V}(M)) \times \Gamma^{2}(T^{*}(M) \otimes V(M)) \times \Gamma^{2} \Big(\bigwedge^{2} T^{*}(M) \otimes \big((T^{*}(M) \otimes T(M)) \times (V^{*}(M) \otimes V(M)) \big) \Big),$$

we can integrate the volume form field $\mathbf{dL}(v, w, r)$ all over M, if it is integrable. If M is compact, every continuous volume form field is integrable on M.

Definition 3. Let M be a compact base manifold, and dL a Lagrange form. Then the action functional defined by the Lagrange form is

$$S: \Gamma^{3}(\breve{V}(M)) \times \breve{D}^{3}(T(M), V(M)) \to \mathbb{R}, (v, \nabla) \mapsto S_{v, \nabla} := \int_{M} \mathbf{dL}(v, \nabla v, F_{\nabla}),$$

where F_{∇} is the curvature tensor of the covariant derivation ∇ .

Unfortunately, if M is noncompact, we cannot proceed with the straightforwardness as in the compact case. If we would like to extend our formalism to a noncompact base manifold, we should proceed otherways. A possible way to realize the action functional, over a noncompact base manifold, could be to define it as a real valued Radon measure on the subsets of the manifold. If we follow this idea, we can make the following definition.

Definition 4. Let the base manifold M be noncompact. If

$$(v, \nabla) \in \Gamma^3(\breve{V}(M)) \times \breve{D}^3(T(M), V(M))$$

and K is a compact set in M, then let us define $S_{v,\nabla}(K) := \int_{K} \mathbf{dL}(v, \nabla v, F_{\nabla})$. The map $K \mapsto S_{v,\nabla}(K)$ uniquely extends to a real valued Radon measure on the Baire quasi- σ -ring of M. Let $\operatorname{Rad}(M, \mathbb{R})$ be the real vector space of the real valued Radon measures on the Baire quasi- σ -ring of M. Then, the action functional is defined as the Radon measure valued map

$$S: \Gamma^{3}(\check{V}(M)) \times \check{D}^{3}(T(M), V(M)) \to \operatorname{Rad}(M, \mathbb{R}), (v, \nabla) \mapsto S_{v, \nabla}.$$

Of course, this definition is also meaningful for the compact case, and the action in the compact case can be expressed as $(v, \nabla) \mapsto S_{v, \nabla}(M)$.

3 The field equations as Euler-Lagrange equations

3.1 Natural distribution topologies on the sections of a vector bundle

Let us take a real $C^{\check{k}}$ vector bundle W(N) over the real C^k manifold N ($\check{k} \in \{0, \ldots, k\}$). Then we can define C^l norm fields ($l \in \{0, \ldots, \check{k}\}$) on $\Gamma^l(W(N))$. Let us take a map

$$\|\cdot\|:\Gamma^{l}(W(N))\to\Gamma^{0}(F(N)),s\mapsto\|s\|,$$

⁸The definitions used by classic literature differ here: as pointed out earlier, there are three kinds of definitions, and all the three approaches have certain problems. The alternative definition, which we give, mostly resembles to the second approach, as we define the action functional (in the case of a noncompact base manifold) as a real valued Radon measure on the Baire quasi- σ -ring of the base manifold (i.e. on the quasi- σ -ring, generated by the compact subsets of the base manifold).

which is pointwise, that is

$$\forall p \in N : \forall s, s' \in \Gamma^l(W(N)) : s'(p) = s(p) \Rightarrow ||s'||(p) = ||s||(p)$$

holds (this means, that it can be viewed as a C^0 fiber bundle homomorphism). If for every $p \in N$ the map $\|\cdot\|_p : W_p(N) \to \mathbb{R}$, naturally defined by the restriction of $\|\cdot\|$, is a norm, then we call $\|\cdot\|$ a C^l norm field. It is a fact that every $C^{\check{k}}$ vector bundle over a C^k manifold admits C^l norm fields: by the paracompactness of manifolds⁹, there are C^l Riemann metric tensor fields on the given vector bundle, and they naturally give rise to C^l norm fields by taking the pointwise norms generated by them (but not all C^l norm fields can be formulated in this way).

Lemma 5. Let N, W(N) be as above. If $\|\cdot\|$ and $\|\cdot\|'$ are C^l norm fields on W(N), then there exists a positive $c \in \Gamma^k(F(N))$, such that $\|\cdot\|' \leq c \|\cdot\|$.

Proof The proof is based on the paracompactness of the differentiable manifolds and on the equivalence of norms on a finite dimensional vector space.

Let us take a locally finite atlas $((U_i, \varphi_i, f_i))_{i \in I}$ of N with partition of unity, such that every U_i $(i \in I)$ has compact closure. Let us denote the closure of a set U with \overline{U} . Let n be the dimension of the fibers of W(N). Let us fix a trivialization $(e_{i,j})_{j \in \{1,...,n\}}$ of W(N) over each chart (U_i, φ_i) $(i \in I)$.

As a consequence of the equivalence of norms on a finite dimensional vector space, for every $p \in N$ there is a positive number c_p , such that $\|\cdot\|'_p \leq c_p \|\cdot\|_p$. Furthermore, c_p can be chosen

to be $\sup_{s_p \in W_p(N) \setminus \{0_p\}} \frac{\|s_p\|'_p}{\|s_p\|_p}$.

It is easily checked, that the equality

$$\sup_{p \in \overline{U_i}} \sup_{s_p \in W_p(N) \setminus \{0_p\}} \frac{\|s_p\|_p'}{\|s_p\|_p} = \sup_{p \in \overline{U_i}} \sup_{S \in \mathbb{R}^n, |S|=1} \frac{\|\sum_{j=1}^n S_j e_{i,j}\|'(p)}{\|\sum_{j=1}^n S_j e_{i,j}\|(p)}$$

holds $(i \in I)$. The rightside is a finite positive number, because it can be viewed as the maximum of a positive valued continuous function over the compact manifold $\overline{U_i} \times \mathbb{S}^{n-1}$ (\mathbb{S}^{n-1} is the n-1dimensional unit sphere). Let us denote this positive number by c_i . Then: $\|\cdot\|' \leq c_i \|\cdot\|$ holds over $\overline{U_i}$ $(i \in I)$.

As a consequence of the previous inequality: $f_i \| \cdot \|' \leq c_i f_i \| \cdot \|$ holds all over the manifold N for each $i \in I$, as a consequence of the fact that f_i is nonnegative and $\operatorname{supp}(f_i) \subset U_i$. The $\operatorname{sum} \sum_{i \in I} c_i f_i$ (which has only finite nonzero terms in a small neighborhood of every point, as a consequence of the local finiteness of the atlas) is a positive valued C^k function, furthermore the sum $\sum_{i \in I} f_i$ (which also has only finite nonzero terms in a small neighborhood of every point)

is 1 by definition. Therefore $\|\cdot\|' \leq \left(\sum_{i \in I} c_i f_i\right) \|\cdot\|$ holds, so the lemma is proved.

We call this property the *equivalence* of C^l norm fields, in the analogy of the equivalence of norms on a finite dimensional vector space.

⁹A manifold is paracompact as a consequence of its definition.

Let us take a $C^{\tilde{k}}$ -type covariant derivation ∇ from $D^{\tilde{k}}(T(N)) \times D^{\tilde{k}}(W(N))$ ($\tilde{k} \in \{0, \dots, \check{k}\}$). If we take $C^{\tilde{k}-l}$ norm fields

$$\|\cdot\|_{l}: \Gamma^{\tilde{k}-l}((\overset{l}{\otimes}T^{*}(N))\otimes W(N)) \to \Gamma^{0}(F(M))$$

for each $l \in \{0, \ldots, \tilde{k}\}$, then we can formulate the quantity

$$\sum_{l=0}^{\tilde{k}} \|\nabla^{(l)} \cdot\|_l : \Gamma^{\tilde{k}}(W(N)) \to \Gamma^0(F(M)), s \mapsto \sum_{l=0}^{\tilde{k}} \|\nabla^{(l)}s\|_l$$

(this is not a norm field, because it is not pointwise, but it is a similar quantity).

Lemma 6. If we choose norm fields $(\|\cdot\|_l)_{l \in \{0,...,\tilde{k}\}}$, and two $C^{\tilde{k}}$ covariant derivations ∇ and ∇' as above, then there exists a positive C^k function c over N, such that $\sum_{l=0}^{\tilde{k}} \|\nabla'^{(l)}\cdot\|_l \le c \sum_{l=0}^{\tilde{k}} \|\nabla^{(l)}\cdot\|_l$.

Proof This is a consequence of the following facts:

- 1. the covariant derivation ∇' can be expressed as the sum of ∇ and a $C^{\tilde{k}-1}$ diagonal Christoffel tensor field,
- 2. the triangle inequality of the norms,
- 3. the composition of a norm with a linear map is a semi-norm,
- 4. the sum of a norm and a semi-norm is a norm,
- 5. Lemma 5.

Corollary 7. If we take norm fields $(\|\cdot\|_l)_{l \in \{0,...,\tilde{k}\}}$, $(\|\cdot\|'_l)_{l \in \{0,...,\tilde{k}\}}$ and covariant derivations ∇ and ∇' as above, then there exists a positive C^k function c over N, such that

$$\sum_{l=0}^{\tilde{k}} \| \nabla'^{(l)} \cdot \|_{l}' \le c \sum_{l=0}^{\tilde{k}} \| \nabla^{(l)} \cdot \|_{l}.$$

Proof This is a consequence of Lemma 6 and Lemma 5.

Corollary 7 lets us define the notions of distribution topologies on the vector space of sections of vector bundles.

Definition 8. Let $\mathcal{E}^{\tilde{k}}(W(N)) := \Gamma^{\tilde{k}}(W(N))$ with the natural real vector space structure. Let us choose a class of norm fields $(\|\cdot\|_l)_{l\in\{0,\dots,\tilde{k}\}}$ and a covariant derivation ∇ as before. Let $\psi, \varphi_n \in \mathcal{E}^{\tilde{k}}(W(N))$ $(n \in \mathbb{N})$, then the sequence $n \mapsto \varphi_n$ converges to ψ in $\mathcal{E}^{\tilde{k}}$ -sense, if and only if the function $\sum_{l=0}^{k} \|\nabla^{(l)}(\psi - \varphi_n)\|_l$ converges to zero uniformly on every compact set. This notion uniquely characterizes a topology on $\mathcal{E}^{\tilde{k}}(W(N))$, which is called the $\mathcal{E}^{\tilde{k}}$ -topology. Note, that as a consequence of Corollary 7, this notion is independent of the chosen norm fields and covariant derivation.

Definition 9. Let $\mathcal{D}^{\tilde{k}}(W(N))$ be the set of elements of $\Gamma^{\tilde{k}}(W(N))$, which have compact support. $\mathcal{D}^{\tilde{k}}(W(N))$ has a natural real vector space structure. Let us choose a class of norm fields $(\|\cdot\|_l)_{l\in\{0,\ldots,\tilde{k}\}}$ and a covariant derivation ∇ as before. Let $\psi, \varphi_n \in \mathcal{D}^{\tilde{k}}(W(N))$ $(n \in \mathbb{N})$, then the sequence $n \mapsto \varphi_n$ converges to ψ in $\mathcal{D}^{\tilde{k}}$ -sense, if and only if there exists a compact set K, such that $\forall n \in \mathbb{N}$: $\operatorname{supp}(\psi - \varphi_n) \subset K$ and the function $\sum_{l=0}^{\tilde{k}} \|\nabla^{(l)}(\psi - \varphi_n)\|_l$ uniformly converges to zero. This notion uniquely characterizes a topology on $\mathcal{D}^{\tilde{k}}(W(N))$, which is called the $\mathcal{D}^{\tilde{k}}$ -topology. Note, that as a consequence of Corollary 7, this notion is independent of the chosen norm fields and covariant derivation.

Definition 10. Let N be compact. Then $\mathcal{E}^{\tilde{k}}(W(N)) = \mathcal{D}^{\tilde{k}}(W(N))$, and the $\mathcal{E}^{\tilde{k}}$ and $\mathcal{D}^{\tilde{k}}$ topologies are the same. Furthermore, if we choose a class of norm fields $(\|\cdot\|_l)_{l\in\{0,...,\tilde{k}\}}$ and a
covariant derivation ∇ as before, then the quantity $\sup_{N} \sum_{l=0}^{\tilde{k}} \|\nabla^{(l)} \cdot\|_l$ is always finite, and this is
a complete norm on $\mathcal{E}^{\tilde{k}}(W(N))$. If we take an other class of norm fields $(\|\cdot\|_l)_{l\in\{0,...,\tilde{k}\}}$ and a
covariant derivation ∇' as before, then there is a positive number c, such that

$$\sup_{N} \sum_{l=0}^{\tilde{k}} \| \nabla'^{(l)} \cdot \|_{l}' \le c \sup_{N} \sum_{l=0}^{\tilde{k}} \| \nabla^{(l)} \cdot \|_{l},$$

that is the two norms are equivalent. Let us call them $C^{\tilde{k}}$ -norms of $\mathcal{E}^{\tilde{k}}(W(N))$. It is easily seen, that the $\mathcal{E}^{\tilde{k}}$ -topology is the same as the topology generated by any $C^{\tilde{k}}$ -norm on $\mathcal{E}^{\tilde{k}}(W(N))$. These are the consequences of Corollary 7 and of the fact, that the continuous real valued functions over a compact manifold are bounded.

3.2 The derivative of action functional

Let us take $(v, \nabla), (v', \nabla') \in \Gamma^3(\breve{V}(M)) \times \breve{D}^3(T(M), V(M))$, then we can express the primed quantities as $v' = v + \delta v, \ \nabla' = \nabla + \delta C$, where the vector field δv is C^3 and the diagonal Christoffel tensor field δC is C^2 .

Theorem 11. Let (v, ∇) and $(\delta v, \delta C)$ be the above quantities, and K a compact subset of M. Then

$$S_{v+\delta v,\nabla+\delta C}(K) = S_{v,\nabla}(K) +$$

$$\begin{split} &\int_{K} \left(D_{1} \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta v + D_{2} \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) (\nabla \delta v + \delta C v + \delta C \delta v) + D_{3} \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) (F_{\nabla + \delta C} - F_{\nabla}) \right) \\ &+ \int_{K} \frac{1}{2} \left[\delta v \quad (\nabla \delta v + \delta C v + \delta C \delta v) \quad (F_{\nabla + \delta C} - F_{\nabla}) \right] \cdot \left[D^{(2)} \mathbf{d} \mathbf{L} \right] (v + \delta v', \nabla v + \delta A', F_{\nabla} + \delta F') \cdot \\ & \left[\begin{pmatrix} \delta v \\ (\nabla \delta v + \delta C v + \delta C \delta v) \\ (F_{\nabla + \delta C} - F_{\nabla}) \end{pmatrix} \right] \end{split}$$

for some sections $\delta v'$, $\delta A'$, $\delta F'$, where for each $p \in K$ there exists a number $c_p \in]0, 1[$, such that $\delta v'_p = c_p \delta v_p$, $\delta A'_p = c_p (\nabla \delta v|_p + \delta C v|_p + \delta C \delta v|_p)$ and $\delta F'_p = c_p (F_{\nabla + \delta C}|_p - F_{\nabla}|_p)$. Here the action

of a $T^*(M) \otimes ((T(M) \otimes T^*(M)) \times (V(M) \otimes V^*(M)))$ type tensor field δC on a V(M) type tensor field δv is defined by the contraction of the projection to the $T^*(M) \otimes V(M) \otimes V^*(M)$ component of δC by δv .

Proof This is a simple consequence of the Taylor formula for one dimensional vector space valued C^2 functions, applied to the Lagrange form, in every point of K:

$$\mathbf{dL}(v+\delta v, (\nabla+\delta C)(v+\delta v), F_{\nabla+\delta C}) = \mathbf{dL}(v, \nabla v, F_{\nabla}) +$$

$$\begin{pmatrix} D_{1}\mathbf{d}\mathbf{L}(v,\nabla v,F_{\nabla})\delta v + D_{2}\mathbf{d}\mathbf{L}(v,\nabla v,F_{\nabla})(\nabla\delta v + \delta C v + \delta C \delta v) + D_{3}\mathbf{d}\mathbf{L}(v,\nabla v,F_{\nabla})(F_{\nabla+\delta C} - F_{\nabla}) \end{pmatrix} + \\ \frac{1}{2} \begin{bmatrix} \delta v & (\nabla\delta v + \delta C v + \delta C \delta v) & (F_{\nabla+\delta C} - F_{\nabla}) \end{bmatrix} \cdot \begin{bmatrix} D^{(2)}\mathbf{d}\mathbf{L} \end{bmatrix} (v + \delta v',\nabla v + \delta A',F_{\nabla} + \delta F') \cdot \\ \begin{bmatrix} \delta v \\ (\nabla\delta v + \delta C v + \delta C \delta v) \\ (F_{\nabla+\delta C} - F_{\nabla}) \end{bmatrix}$$

holds, with the notations in the statement of the theorem.

We have to show only that the term under the last integral (the term with the second derivatives) is integrable. This follows from the simple fact: we can express it via the terms with lower derivatives by using the Taylor formula. The terms with lower derivatives are C^1 by Remark 2. Therefore, the term with the second derivatives is C^1 , although the function $p \mapsto c_p$ need not be even measurable. So, the last term is integrable on K.

3.2.1 The case of compact base manifold

Our intention is to interpret the linear term of the expression in *Theorem 11* as a kind of derivative of the action functional. If M is compact, this can be realized quite straightforward, by using *Definition 10*.

Let us assume that M is compact, without boundary. Let us take the standard topology on \mathbb{R} . Let us fix a C^3 -norm on $\mathcal{E}^3(V(M))$ and a C^2 -norm on

$$\mathcal{E}^{2}\Big(T^{*}(M)\otimes\big((T(M)\otimes T^{*}(M))\times(V(M)\otimes V^{*}(M))\big)\Big).$$

Then we can fix a power q (which is a real number greater or equal to 1, or infinity) and take the L^q product norm of these norms on the product space. With this, the space

$$\Gamma^{3}(V(M)) \times \left(D^{3}(T(M)) \times D^{3}(V(M)) \right)$$

forms an affine space over the normed space

$$\mathcal{E}^{3}(V(M)) \times \mathcal{E}^{2}\Big(T^{*}(M) \otimes \big((T(M) \otimes T^{*}(M)) \times (V(M) \otimes V^{*}(M))\big)\Big)$$

In this sense, we can take the derivative of the action functional. The derivative is independent of the chosen C^3 and C^2 -norms, and of the power q (that is, of the way of forming of the product norm), because the notion of derivative depends only on the equivalence class of norms.¹⁰

¹⁰If we choose other C^3 and C^2 -norms on the previous spaces, and take an other power q for forming an L^q product norm, then we get a new norm, which is equivalent to the previous one.

Theorem 12. The action functional S is continuously differentiable, and its derivative at given (v, ∇) is the continuous linear map $(\delta v, \delta C) \mapsto$

$$\int_{M} \Big(D_1 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta v + D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) (\nabla \delta v + \delta C v) + D_3 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) 2 \nabla \wedge \delta C \Big),$$

where the wedge in the expression $\nabla \wedge \delta C$ means antisymmetrization in the $T^*(M)$ variable of ∇ and the first $T^*(M)$ variable of δC in the expression $\nabla \delta C$.

Proof Note that this expression is the linear term from the expression in Theorem 11. First, we have to show that this linear map is a continuous linear map, and that the remaining bilinear term from Theorem 11 is an ordo function. By the differentiability properties of dL, and the compactness of M, these facts are direct consequences of Lebesgue theorem. Finally, the derivative function of S is continuous, for the same reason.

Theorem 13. The derivative of the action functional S can be expressed at given (v, ∇) as the continuous linear map

$$(\delta v, \delta C) \mapsto \int_{M} \Big(D_1 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta v - (\nabla \cdot D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla})) \delta v - (\operatorname{Tr} T_{\nabla} \cdot D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla})) \delta v \Big) + \int_{M} \Big(D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta C v - 2(\nabla \cdot \widehat{D_3 \mathbf{d} \mathbf{L}}(v, \nabla v, F_{\nabla})) \delta C - 2(\operatorname{Tr} T_{\nabla} \cdot \widehat{D_3 \mathbf{d} \mathbf{L}}(v, \nabla v, F_{\nabla})) \delta C \Big).$$

Here T_{∇} is the torsion tensor of ∇ , $\operatorname{Tr} T_{\nabla}$ denotes the contraction of the second $T^*(M)$ and the T(M) variable of T_{∇} . The hat in the $\widehat{D_3 \mathbf{dL}}$ expression means antisymmetrization in the first two T(M) variables. Finally, \cdot means contraction of the $T^*(M)$ variable of ∇ and the first T(M) variable of the tensor quantity after it, or the contraction of the $T^*(M)$ variable of $\operatorname{Tr} T_{\nabla}$ and the first T(M) variable of the tensor quantity after it, respectively.

Let us call the equality of the above map and the derivative of S the Euler-Lagrange relation, and let us call the above map the Euler-Lagrange map. \Box

Proof We simply make transformations of the expression in Theorem 12.

The term $D_3 \mathbf{dL}(v, \nabla v, F_{\nabla}) 2\nabla \wedge \delta C$ is equal to $2D_3 \mathbf{dL}(v, \nabla v, F_{\nabla}) \nabla \delta C$, this is easily seen for example by using Penrose abstract indices:

$$D_3^{ab} \mathbf{dL}(v, \nabla v, F_{\nabla}) (\nabla_a \delta C_b - \nabla_b \delta C_a) = \left(D_3^{ab} \mathbf{dL} - D_3^{ba} \mathbf{dL} \right) (v, \nabla v, F_{\nabla}) \nabla_a \delta C_b,$$

where the indices a, b indicate the T(M) or $T^*(M)$ variables in question.

By the Leibniz rule

$$D_{2}\mathbf{dL}(v,\nabla v,F_{\nabla})\nabla\delta v + 2\widehat{D_{3}\mathbf{dL}}(v,\nabla v,F_{\nabla})\nabla\delta C =$$

$$\nabla \cdot \left(D_{2}(\mathbf{dL}(v,\nabla v,F_{\nabla})\delta v) + \nabla \cdot \left(2\widehat{D_{3}\mathbf{dL}}(v,\nabla v,F_{\nabla})\delta C\right) - \left(\nabla \cdot D_{2}(\mathbf{dL}(v,\nabla v,F_{\nabla})\right)\delta v - \left(\nabla \cdot 2\widehat{D_{3}\mathbf{dL}}(v,\nabla v,F_{\nabla})\right)\delta C$$

is true.

The sum of the first two terms on the rightside of the equation can be written in the form $\nabla \cdot \mathbf{dA}$, where $\mathbf{dA} \in \Gamma^1(T(M) \otimes \bigwedge^m T^*(M))$ (that is \mathbf{dA} is a C^1 volume form valued vector field). Let us take an other covariant derivation $\tilde{\nabla}$ on the tensor bundles of T(M), which is the Levi-Civita covariant derivation of some semi-Riemannian metric tensor field \tilde{g} over M. If $C \in \Gamma^3(T^*(M) \otimes T(M) \otimes T^*(M))$ is the Christoffel tensor field on T(M) of ∇ relative to $\tilde{\nabla}$, then one can figure out the fact, that $\nabla \cdot \mathbf{dA} = \tilde{\nabla} \cdot \mathbf{dA} + (\operatorname{Tr}_1 C - Tr_2 C) \cdot \mathbf{dA}$, where $\operatorname{Tr}_1 C$ denotes the contraction of the first $T^*(M)$ and the T(M) variable of C, and $\operatorname{Tr}_2 C$ denotes the contraction of the second $T^*(M)$ and the T(M) variable of C. It is easier to follow the previous statement in Penrose abstract indices: $\nabla_a \mathbf{dA}^a = \tilde{\nabla}_a \mathbf{dA}^a + (C_{ba}^b - C_{ab}^b)\mathbf{dA}^a$, because $\nabla_a t^b = \tilde{\nabla}_a t^b + C_{ac}^b t^c$ is valid for a tangent vector field t, and $\nabla_a \mathbf{dv} = \nabla_a \mathbf{dv} - C_{ab}^b \mathbf{dv}$ is true for a volume form field \mathbf{dv} . As $\tilde{\nabla}$ is a Levi-Civita covariant derivation, the $(\operatorname{Tr}_1 C - Tr_2 C)$ quantity corresponds to $-\operatorname{Tr} T_{\nabla}$, because the torsion of $\tilde{\nabla}$ vanishes by definition. Thus, we can write $\nabla \cdot \mathbf{dA} = \tilde{\nabla} \cdot \mathbf{dA} - \operatorname{Tr} T_{\nabla} \cdot \mathbf{dA}$.

The term with the torsion corresponds to the term with the torsion in the statement of the theorem. To prove the theorem, we only have to show, that the integral of $\tilde{\nabla} \cdot \mathbf{dA}$ is zero.

Let us use the fact that M is orientable: there exists a nowhere zero C^3 volume form field \mathbf{dv} . As the vector space of the volume forms, at a given point, is one dimensional, then we can uniquely define a nowhere zero section \mathbf{dp} of the dual volume form bundle, such that at every point, \mathbf{dp} maps \mathbf{dv} into 1. By using coordinate charts, it can be seen, that \mathbf{dp} is also C^3 . Then by contracting the quantity $\mathbf{dp} \otimes \mathbf{dA}$ in the volume form and dual volume form variables, one can define a C^1 vector field (as \mathbf{dA} is C^1). Let us denote this by \mathbf{dA}/\mathbf{dv} . With the introduced notation, one has $\mathbf{dA} = \mathbf{dv} \otimes (\mathbf{dA}/\mathbf{dv})$.

Let $\mathbf{d}\mathbf{v}$ be one of the two canonical volume form fields associated to \tilde{g} . Then $\tilde{\nabla}\mathbf{d}\mathbf{v} = 0$ holds, which implies by the Leibniz rule: $\tilde{\nabla} \cdot \mathbf{d}\mathbf{A} = \mathbf{d}\mathbf{v}\,\tilde{\nabla} \cdot (\mathbf{d}\mathbf{A}/\mathbf{d}\mathbf{v})$. It is a theorem, that if Xis a tangent vector field, then $\mathbf{d}\mathbf{v}\,\tilde{\nabla}\cdot X = m\,\mathbf{d}(X.\mathbf{d}\mathbf{v})$, where m is the dimension of M, \mathbf{d} means the exterior derivation, and the dot . means contraction of X with the first $T^*(M)$ variable of $\mathbf{d}\mathbf{v}$ (see for example [5]). By using the observations in the previous paragraph, we can state that $(\mathbf{d}\mathbf{A}/\mathbf{d}\mathbf{v}).\mathbf{d}\mathbf{v} = \operatorname{Tr} \mathbf{d}\mathbf{A}$, where Tr means the contraction of the T(M) variable of $\mathbf{d}\mathbf{A}$ with the first $T^*(M)$ variable of $\mathbf{d}\mathbf{A}$. We get, that the expression $\mathbf{d}\mathbf{v}\,\tilde{\nabla}\cdot(\mathbf{d}\mathbf{A}/\mathbf{d}\mathbf{v})$ is independent of the choice of the semi-Riemannian metric tensor field \tilde{g} , and it is equal to $\mathbf{d}(m \operatorname{Tr} \mathbf{d}\mathbf{A})$, the exterior derivative of the (m-1)-form field $m \operatorname{Tr} \mathbf{d}\mathbf{A}$. The integral of this term vanishes as a consequence of Gauss theorem, because M is a compact manifold without boundary. So the formula, stated in the theorem, is valid.

Remark 14. If M is a compact manifold with boundary, then the presented statements remain true, but the Euler-Lagrange map in Theorem 13 has an extra term, which is a boundary integral, as a consequence of the Gauss theorem. Namely, the derivative of S at given (v, ∇) is the continuous linear map

$$\begin{split} (\delta v, \delta C) &\mapsto \int_{M} \Big(D_1 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta v - (\nabla \cdot D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla})) \delta v - (\operatorname{Tr} T_{\nabla} \cdot D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla})) \delta v \Big) + \\ \int_{M} \Big(D_2 \mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta C v - 2 (\nabla \cdot \widehat{D_3 \mathbf{d} \mathbf{L}}(v, \nabla v, F_{\nabla})) \delta C - 2 (\operatorname{Tr} T_{\nabla} \cdot \widehat{D_3 \mathbf{d} \mathbf{L}}(v, \nabla v, F_{\nabla})) \delta C \Big) + \\ & m \int_{\partial M} \operatorname{Tr} \Big(D_2 (\mathbf{d} \mathbf{L}(v, \nabla v, F_{\nabla}) \delta v + 2 \widehat{D_3 \mathbf{d} \mathbf{L}}(v, \nabla v, F_{\nabla}) \delta C \Big), \end{split}$$

where ∂M is the boundary of M.

As a summary, we can define a classical field theory over a compact base manifold M (with or without boundary) as a quartet $(M, V(M), \mathbf{dL}, S)$, where V(M) is a vector bundle as in the text, \mathbf{dL} is a Lagrange form, and S is the action functional, defined by \mathbf{dL} . The field equation is the equation

$$\left((v,\nabla)\in\Gamma^3(\breve{V}(M))\times\breve{D}^3(T(M),V(M))\right)\ ?\quad DS(v,\nabla)=0,$$

where DS denotes the derivative of S. Let us call DS the Euler-Lagrange functional.

3.2.2 The case of noncompact base manifold

If the base manifold is noncompact, the vector spaces of sections of a vector bundle do not have natural normed space structure, they only have natural \mathcal{E} or \mathcal{D} distribution topologies.

Let us take a map $Q : R \to S$, where R and S are some spaces. To be able to define the classical (Fréchet) notion of derivative of Q, the space R has to be a normed affine space, and S has to be a topological vector space (in order to be able to define the notion of ordo functions).

In the case of the action functional, when the base manifold is noncompact, the first space is the space $\Gamma^3(V(M)) \times (D^3(T(M)) \times D^3(V(M)))$, which forms an affine space over the topological vector space $\mathcal{E}^3(V(M)) \times \mathcal{E}^2(T^*(M) \otimes ((T(M) \otimes T^*(M)) \times (V(M) \otimes V^*(M)))))$. The second space is the space of real valued Radon measures $\operatorname{Rad}(M, \mathbb{R})$, which is a vector space, and it possesses a natural topology, uniquely characterized by the following notion of limes: a sequence of Radon measures converges to a Radon measure, if both evaluated on any fixed compact set of M, the sequence of values (real numbers) converges to the value of the given Radon measure (real number). (This is the *pointwise*, or the *setwise* topology on the Radon measures.)

As we see, when M is noncompact, the notion of the derivative of S cannot be defined: the obstruction is that $\Gamma^3(V(M)) \times (D^3(T(M)) \times D^3V(M)))$ only has a natural topological affine space structure, instead of a natural normed affine space structure. Thus, if we want to proceed in the noncompact case, and want to define a similar quantity to an Euler-Lagrange functional, we can not interpret it as a (Fréchet) derivative.

There are known constructions, which are based on a formulation popular in physics literature, even in mathematical physics literature (see e.g. [5], [8]). It defines $(v, \nabla) \mapsto DS_{v,\nabla}(K)$ by using one-parameter families of field configurations, which are fixed on the boundary of a fixed compact set K with smooth boundary. (We will refer to these formulations as *one-parameter* family formulations.) To define the field equations, they take a covering $(K_i)_{i\in I}$ of M with such compact sets, and on every set they require $DS(K_i) = 0$ $(i \in I)$. It can be proved, that $DS(K_i) = 0$ means Euler-Lagrange equations over the interior of the given compact set K_i $(i \in I)$, so after all, over the whole spacetime manifold M. This statement is true, but we have one more constraint: the field values are fixed on the system of boundaries $(\partial K_i)_{i\in I}$.

As one can see, the one-parameter family construction is quite cumbersome. Furthermore, it is not constructive in the following sense. Let us fix a compact set with sooth boundary, and a field configuration on the boundary. If the Euler-Lagrange equations are first order hyperbolic (e.g. when the Dirac equation is part of the field equations), then generally there is no such field configuration on the compact set, which satisfies the field equations and has the (arbitrarily) chosen boundary values. Thus, one can not generate solutions inside a compact set by specifying (arbitrary) boundary values.

There are constructions known in physics literature, which are defined otherways. We shall refer to these as *time-slice constructions*. These assume a cylindric base manifold, i.e. a manifold diffeomorphic to $\mathbb{R} \times C$ for some manifold C (which will be referred as space or time-slice). The action is defined as the integral of the Lagrange form on the domain between two specific time-slice. Certain spatial fall-off properties have to be introduced in order to be able to define the action functional, if C is not compact. The Euler-Lagrange functional is then defined as the derivative of the action with respect to appropriate C^k supremum norms (for some $k \in \mathbb{N}$), similar to the case of compact base manifold. The problem is: how to formalize the spatial regularity conditions. In the literature this problem is carefully overlooked, if possible. The most self-suggesting solution seems to be to introduce a global coordinate system on C (this is, of course, not always possible), and treat the fall-off properties with respect to the coordinates. This method would be quite inelegant (as it refers to global coordinate chart), furthermore it would highly depend on this preferred coordinate system.

For the above problem *Philip E. Parker* suggested us a partial solution, which avoids coordinate systems. He drew our attention to his work [1], which partly deals with a problem of fall-off properties. In his paper, he uses the topological approach to infinities of manifolds: the set of *ends* of the manifold C can be defined as $E(C) := \liminf_{\substack{K \subset C \\ \text{compact}}} \pi_0(C \setminus K)$, where $\pi_0(C \setminus K)$

means the set of the connected components of $C \setminus K$, and liminv is the so called inverse limes, known in topology. An end represents an infinity in the topological sense. Then, he is able to define when two Riemannian metric tensor fields (of some vector bundle) falls off at a given infinity in the same way (notion of order relatedness). This notion provides an equivalence class concept between Riemannian metric tensor fields. Given such an equivalence class of Riemannian metric tensor fields, one can define the notion of *rapidly decreasing* field configurations, which can be used to introduce fall-off properties. However, as indicated, this concept highly depends on the used metric tensor field equivalence class, the (physical) meaning of which is quite unclear (just as in the case of preferring a global coordinate system on C). Furthermore, the method would also highly depend on the initial splitting of the spacetime manifold into $\mathbb{R} \times C$, which conflicts with the philosophy of the theory of relativity.

4 Discussion

We have seen, that the variational formulation of general relativistic field theories can be defined with a significant mathematical elegance over compact base manifolds (with or without boundary). Over noncompact base manifolds, the variational principles can be defined with a great effort, the known constructions are not elegant at all in mathematical sense, furthermore they do have problems with the interpretation.¹¹

In physics, it is held as a principle, that the equation of motion of fields arise from some Euler-Lagrange equations (that is, as some equation $DS(v, \nabla) = 0$). If we want to preserve this principle, and want to avoid the rather questionable constructions in the noncompact cases at the same time, we can make a choice to solve the problem.

1. We can restrict the spacetime models to compact orientable cases.

¹¹The problem of *non-constructiveness* in the case of one-parameter family constructions, furthermore the problem of *spacetime splitting* and the *metric tensor field equivalence class dependence of fall-off properties* in the case of time-slice constructions.

2. We do not interpret the base manifold as the spacetime manifold itself, but as a kind of compactification of it.

The first case is unacceptable: it is a theorem, that every compact spacetime model admits closed timelike curves. So, a compact spacetime model, arising from any kind of formulation, cannot be considered physically realistic.

The other case does not have physical obstructions, and has a certain mathematical elegance. But then, the question arises: if we do not interpret the base manifold directly as the spacetime manifold, how do we interpret it?

For this problem, a possible solution is the condition of asymptotic simpleness of a spacetime. (See e.g. [5], [8].) If this condition holds, then one can define the notion of conformal infinities of the spacetime and the conformal compactification of the spacetime, which will be a compact manifold with boundary.

From the above argument, it is likely to consider only the compact case of a base manifold (with boundary), and interpret it as the conformal compactification of the spacetime manifold.

Theorem 15. Let the base manifold M be compact with boundary. Let

 $(v, \nabla) \in \Gamma^3(\breve{V}(M)) \times \breve{D}^3(T(M), V(M)).$

Then the condition $DS(v, \nabla) = 0$ is equivalent to the followings:

1. the Euler-Lagrange equations, that is the equations

$$D_1 \mathbf{dL}(v, \nabla v, F_{\nabla}) - (\nabla \cdot D_2 \mathbf{dL}(v, \nabla v, F_{\nabla})) - (\operatorname{Tr} T_{\nabla} \cdot D_2 \mathbf{dL}(v, \nabla v, F_{\nabla})) = 0,$$

$$D_2 \mathbf{dL}(v, \nabla v, F_{\nabla})(\cdot)v - 2(\nabla \cdot \widehat{D_3 \mathbf{dL}}(v, \nabla v, F_{\nabla}))(\cdot) - 2(\operatorname{Tr} T_{\nabla} \cdot \widehat{D_3 \mathbf{dL}}(v, \nabla v, F_{\nabla}))(\cdot) = 0$$

are satisfied on the interior of M, and

2. the boundary constraints, that is the equations

$$\operatorname{Tr}(D_2(\mathbf{dL}(v, \nabla v, F_{\nabla})) = 0,$$
$$\operatorname{Tr}(2\widehat{D_3\mathbf{dL}}(v, \nabla v, F_{\nabla}))(\cdot) = 0$$

are satisfied on the boundary of M.

Proof Let us take such sections $(\delta v, \delta C)$, that their support is in the interior of M. Then, the boundary term is zero in the Euler-Lagrange relation in Remark 14. By the Lagrange lemma, condition 1 is implied.

We know now, that condition 1 holds. This means that the non-boundary term in the Euler-Lagrange relation in Remark 14 is zero. Now taking any sections $(\delta v, \delta C)$, condition 2 is implied, by using the Lagrange lemma on the boundary of M.

The question arises: what do the boundary conditions and the boundary of the base manifold mean? In the next section, we shall investigate the physical meaning of the boundary conditions on the example of empty general relativistic spacetime: we shall show that the boundary represents the conformal boundary (conformal infinity) of the arising spacetime model.

5 Boundary as conformal infinity: the example of empty general relativistic spacetime

Let the base manifold M be 4 dimensional, and let us require that M admits C^3 semi-Riemann metric tensor fields with Lorentz signature (this is known to hold if and only if there exists a nowhere zero C^3 tangent vector field on M).

Let us take the vector bundle $V(M) := F(M) \times \bigvee^2 T^*(M)$ (\vee means symmetrized tensor product). We define the sub fiber bundle $\check{V}(M)$ by the restriction of the fibers of V(M) in the following way: for each point $p \in M$ the fiber is restricted to $\mathbb{R} \times L_p(M)$, where $L_p(M)$ denotes the subset of semi-Riemannian metric tensors with Lorentz signature in $\bigvee^2 T_p^*(M)$. It can be easily shown, that $\check{V}(M)$ is such a sub fiber bundle of the vector bundle V(M), as required in the text.

Let us take the sub affine space of $D^3(T(M)) \times D^3(V(M))$, which has the following property: the sub affine space should consist of those pairs (∇, ∇') , where the covariant derivation ∇' over V(M) corresponds to the covariant derivation obtained by the unique extension of ∇ to $F(M) \times \bigvee^2 T^*(M)$, by using *Remark 1*. This sub affine space can be naturally identified with $D^3(T(M))$, therefore we can define a covariant derivation from this sub affine space to be torsion-free if and only if the corresponding covariant derivation from $D^3(T(M))$ is torsionfree. Let $\check{D}^3(T(M), V(M))$ be the sub affine space of torsion-free covariant derivations of the previous sub affine space. It can be easily shown, that this is a closed sub affine space with respect to the topology defined in *Definition 8*.

In this subsection we will apply the usual formalism of Penrose abstract indices, to denote tensor quantities and various contractions of them.

If $g_{ab}(p)$ is a metric tensor with Lorentz signature from $\sqrt[2]{T_p^*}(M)$ $(p \in M)$, then the inverse metric of it (the corresponding Lorentz metric in $\sqrt[2]{T_p}(M)$) will be denoted by $g^{ab}(p)$. Let us take an orientation of M. One of the two associated volume forms to a $g_{ab}(p)$ Lorentz metric (that one, which corresponds to the chosen orientation) will be denoted by $\mathbf{dv}_q(p)$.

If ∇ is a covariant derivation on T(M), the corresponding Riemann-tensor will be denoted by R_{∇} .

With the above notations, let us take the Lagrange form

$$\mathbf{dL} : ((\varphi, g_{ab}), (D\varphi, Dg_{cd}), (R_{efg}{}^h)) \mapsto \mathbf{dv}_g \varphi^2 g^{ik} \delta^j_l R_{ijk}{}^l,$$

which is the abstraction of the Einstein-Hilbert Lagrangian. The field φ will play the role of the geometrized coupling factor to gravity, that is the inverse of the Planck length.

Theorem 16. The Euler-Lagrange equations of the present field theory are

$$2\mathbf{d}\mathbf{v}_g\varphi g^{ac}\delta^b_d(R_\nabla)_{abc}{}^d = 0,$$

$$-\mathbf{d}\mathbf{v}_{g}\varphi^{2}\left(g^{ae}g^{fc}\delta^{b}_{d}(R_{\nabla})_{abc}{}^{d}-\frac{1}{2}g^{ef}g^{ac}\delta^{b}_{d}(R_{\nabla})_{abc}{}^{d}\right)=0,$$

$$-\nabla_{a}\left(\mathbf{d}\mathbf{v}_{g}\varphi^{2}(g^{ac}\delta^{b}_{d}-g^{bc}\delta^{a}_{d})\right)-(T_{\nabla})^{e}_{ae}\left(\mathbf{d}\mathbf{v}_{g}\varphi^{2}(g^{ac}\delta^{b}_{d}-g^{bc}\delta^{a}_{d})\right)=0,$$

which hold in the interior of M.

The boundary constraints are

$$0 = 0,$$

$$0 = 0,$$

$$(\mathbf{d}\mathbf{v}_g)_{afgh}\varphi^2(g^{ac}\delta^b_d - g^{bc}\delta^a_d) = 0,$$

which hold on the boundary of M.

Proof One can get these equations, by simply substituting **dL** into the formulae in Theorem 15, and by using the identities $\frac{\partial \mathbf{dv}_g}{\partial g_{ef}} = \frac{1}{2}g^{ef}\mathbf{dv}_g$ and $\frac{\partial g^{ac}}{\partial g_{ef}} = -\frac{1}{2}(g^{ae}g^{fc} + g^{af}g^{ec})$, which can be derived easily, but also can be found in [5] or [8].

It is easily seen, that the first Euler-Lagrange equation follows from the second one on the domains, where φ is nowhere zero. Furthermore, the first two boundary constraint is trivial. Let us denote the torsion-free part of ∇ with $\tilde{\nabla}$. Then the third Euler-Lagrange equation is equivalent to the equation $-\tilde{\nabla}_a \left(\mathbf{dv}_g \varphi^2 (g^{ac} \delta^b_d - g^{bc} \delta^a_d) \right) = 0.$

Lemma 17. On those open sets, where φ is nowhere zero, the equation

$$-\tilde{\nabla}_a \Big(\mathbf{d} \mathbf{v}_g \varphi^2 (g^{ac} \delta^b_d - g^{bc} \delta^a_d) \Big) = 0$$

is equivalent to the equation $\tilde{\nabla}_a(\varphi^{-2}g^{bc}) = 0.$

Proof The proof will be performed separately in the two implication directions.

 (\Leftarrow) This way is trivial, it can be shown by direct substitution.

(⇒) To prove this way, let us contract the first equation in its b and d indices. We get $-\tilde{\nabla}_a \left(\mathbf{dv}_g \varphi^2 3 g^{ac} \right) = 0$. Therefore, the first equation implies $\tilde{\nabla}_a \left(\mathbf{dv}_g \varphi^2 g^{bc} \right) = 0$. Let us introduce the rescaled metrics $G_{ab} := \varphi^2 g_{ab}$ and $G^{ab} := \varphi^{-2} g^{ab}$. Then the implied equation can be written into the form $\tilde{\nabla}_a \left(\mathbf{dv}_G G^{bc} \right) = 0$. It can be easily seen, for example by using coordinates and the relation $\frac{\partial \mathbf{dv}_g}{\partial g_{ef}} = \frac{1}{2} g^{ef} \mathbf{dv}_g$, that $\tilde{\nabla}_a \mathbf{dv}_G = \frac{1}{2} \mathbf{dv}_G G^{bc} \tilde{\nabla}_a G_{bc}$ for arbitrary covariant derivatives $\tilde{\nabla}$, from which, by the Leibniz rule we infer that $\mathbf{dv}_G (\tilde{\nabla}_a G^{bc} - \frac{1}{2} G^{bc} G_{de} \tilde{\nabla}_a G^{de}) = 0$. We can drop \mathbf{dv}_G from this equation, because it is nowhere zero on the domain in question. Furthermore, by taking its contraction with G_{bc} , one gets $-G_{de} \tilde{\nabla}_a G^{de} = 0$. Therefore, by using this and the previous equation: $\tilde{\nabla}_a G^{bc} = 0$, so finally $\tilde{\nabla}_a (\varphi^{-2} g^{bc}) = 0$ is implied.

We can summarize now the Euler-Lagrange equations: they are equivalent to

$$\tilde{\nabla}_a(\varphi^2 g_{bc}) = 0,$$

$$\varphi^2 \Big((R_{\nabla})_{acb}{}^c - \frac{1}{2} (\varphi^2 g_{ab}) (\varphi^{-2} g^{ef}) (R_{\nabla})_{ecf}{}^c \Big) = 0$$

on those domains of the interior of M, where φ is nowhere zero. From the definition of $\check{V}(M)$ and $\check{D}^3(T(M), V(M))$ we know, that $\varphi^2 g_{ab}$ has Lorentz signature on the above domains, furthermore $\tilde{\nabla} = \nabla$. Thus, the vacuum Einstein equations turned out to be equivalent to the Euler-Lagrange equations on those domains of the interior of M, where the field φ is nowhere zero.

We can summarize the boundary constraints: they are equivalent to $\varphi|_{\partial M} = 0$, which is also equivalent to $(\varphi^2 g_{ab})|_{\partial M} = 0$. The latter means that the boundary of M is the conformal infinity of the arising spacetime model.

The rescaled metric $\varphi^2 g_{ab}$ can be interpreted physically, as the metric, measured in such units, where the coupling factor of gravity (that is, the inverse of the Planck length) is taken to be 1.

6 Concluding remarks

A mathematically precise global approach was presented, to obtain variational formulation of general relativistic classical field theories. According to the authors' information, there is no such formulation, known in literature. For an overview of variational principles over general relativity, see [6]. For further recent works in topic, see [2], [3], [4], [7].

The development of such a formulation was inspired by some problems of the usual approaches, and by possible future application as a tool in the proof of a global existence theorem: as the approach is global, one would simply have to prove critical point theorems on the action functional in order to obtain a theorem on the global existence of solutions. This would be desirable, as the question of global existence of solutions is unsolved in general relativistic field theories, yet.

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