The case of Quantum Gravity with Spontaneous Collapse

Lajos Diósi

Wigner Research Center for Physics, Budapest Eötvös Loránd University, Budapest

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Abstract 2014

Gravity related spontaneous decoherence: from Wheeler-Bekenstein-Hawking to optomechanics (Erice 2015)

The inception of a universal gravity-related irreversibility took place originally in quantum cosmology but it turned out soon that a universal non-unitary dynamics is problematic itself. Independent investigations of the quantum measurement postulate clarified that a non-unitary dynamics is of interest already in the non-relativistic context. An intricate relationship between Newton gravity and quantized bulk matter might result in universal non-relativistic violation of unitarity - also called spontaneous decoherence. The corresponding gravity-related spontaneous decoherence model is now on the verge of detectability in optomechanical experiments. It is also a toy-model of cosmic quantum-gravitational non-unitarity, illuminating that the bottle-neck of quantum-gravity is the quantum measurement postulate instead of quantum cosmology.

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Abstract 2022

When about half a century ago the concept of universal spontaneous collapse of the wave function was concieved it was an attempt to alter standard non-relativistic quantum physics. As such, it was largely ignored by relativistic field theory and quantum-gravity communities. A central motivation of spontaneous collapse community has been to replace the standard collapse-by-measurement that annoved many. For long time it did not annoy the field theory and quantum-gravity communities. Concept of quantum field theory with certain universal irreversibilities had been initiated very long ago by Wheeler, Hawking and a few others independently from the concept of spontaneous collapse. Over the decades the two concepts have come close and support each other.

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Irrev Quantum Gravity/Cosmology at Planck Scale

Heuristic Arguments within Standard Physics

- Wheeler (1955): foamy space-time at Planckian scale no compact dynamical eq.
- Bekenstein (1972): black-holes behave termodynamically

$$S_{BH} = \frac{k_B}{4} \frac{A_{BH}}{A_{PI}}$$

... and even radiate thermally, Hawking (1973)

- Hawking (1983): unitarity is lost due to instantons $\widehat{\rho} \rightarrow \$\widehat{\rho} \neq \widehat{S}\widehat{\rho}\widehat{S}^{\dagger}$
- Banks-Susskind-Peskin (1984): violation of conservations laws $\dot{\hat{\rho}} = -i[\hat{H}, \hat{\rho}] - \int \int [\hat{Q}(x), [\hat{Q}(y), \hat{\rho}]]h(x-y)d^3xd^3y$

 \widehat{Q} is relativistic quantum field, *h* is positive kernel.

Irrev Quantum Mechanics for Massive Objects

Heuristic **modifications** of Standard Physics Purpose: massive Schrodinger Cats $|f_1\rangle + |f_2\rangle$ decay spontaneously

- Karolyhazy (1966): fluctuations of space-time at Planckian scale G-related qualitative eqs.
- GRW (1986): rare spontaneous localizations of constituents G-unrelated exact eqs.
- D. (1986): fluctuations of Newtonian gravitational field

$$\dot{\widehat{\rho}} = -\frac{i}{\hbar}[\widehat{H},\widehat{\rho}] - \frac{G}{2\hbar} \iint [\widehat{f}(x), [\widehat{f}(y),\widehat{\rho}]] \frac{1}{|x-y|} d^3x d^3y$$

 \widehat{f} is non-relativistic quantized mass density field

• Penrose (1996): uncertainty of time-flow

$$\frac{1}{\tau_{decay}} = \frac{G}{\hbar} \iint [f_1(x) - f_2(x)][f_1(y) - f_2(y)] \frac{1}{|x - y|} d^3 x d^3 y$$

 f_1, f_2 mass densities of Cat state

G-related spontaneous decoherence

Particular purpose: $|f_1\rangle + |f_2\rangle$ decay into mixture of $|f_1\rangle$ and $|f_2\rangle$.

Construction of G-related spontaneous decoherence (with one eye on G-related spontaneous collapse):

- formal von Neumann measurements of local mass densities f(x)
- detectors are hidden this time!
- nobody reads out the measurement outcomes

Resulting Master Equation of G-related spontaneous decoherence: $\dot{\hat{\rho}} = -\frac{i}{\hbar}[\hat{H},\hat{\rho}] - \frac{G}{2\hbar}\int[\hat{f}(x),[\hat{f}(y),\hat{\rho}]]\frac{1}{|x-y|}d^{3}xd^{3}y$ $\hat{f} \text{ is non-relativistic quantized mass density field: } \hat{f}(x) = \sum_{n} m_{n}g_{\sigma}(x-\hat{q}_{n}).$

Note: same structure as BSP eq., interpretation is very different.

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Fundamental irreversibility? — Parallel History

	QUANTUM COSMOLOGISTS	SCHRODINGER-CAT KILLERS [†]
1936	Bronstein: ambiguity δg_{ab}	
1950	Wheeler: space-time foam	
1966		Károlyházi: δg_{ab} collapses Ψ
1972	Bekenstein: black hole entropy	
1973	Hawking: black hole radiates	
1983	Hawking: $ ho_f = \$ ho_i eq S ho S^\dagger$	
1984	BanksSusskindPeskin: $T^{ab}_{,b} \neq 0$	
1986		D.: δg_{ab} collapses Ψ , master eq.
		GRW: toy model of Ψ -collapse
1990		GRWP: CSL model of Ψ -collapse
1996	Penrose: δg_{ab} collapses Ψ	
2008	Hogan: holographic noise	
[†] Researchers of Spontaneous Ψ Collapse.		
SchCatKillers: Pearle, D., Bassi's, Tumulka, Tilloy, Bedingham, Laloe,		
COSMOLOGISTS may profit from results of SchCAT KILLERS.		
Some already did, e.g. BL Hu, TP SIngh, Sudarsky, Oppenheim, and		

David Poulin — On Information Loss

Conclusion

- Models of information loss that
 - do not violently break well established principles;
 - are well formulated mathematically; and
 - agree with experiments;

have not been ruled out.

- The secret sauce in our model is violation of causality at microscopic scales.
- Fundamental non-unitary evolution opens up new possibilities for quantum-classical evolution:
 - Further justifies non-unitary evolution since dissipative terms can be controlled by classical gravitational variables: turn on only in extreme conditions.
- To do:
 - Explicitly write rate equation for gravitational field.
 - Work out model details to provide experimental test to refute.

David Poulin — A relativistic Lindblad Eq.

David Poulin — A relativistic Lindblad Eq.

A free field model

A model

- Start with a free scalar theory $H = \frac{1}{2} \int \frac{d^3p}{(2\pi)^3} (\pi^2 + m^2 \phi^2 + (\nabla \phi)^2).$
- Consider positive frequency component of field operators $\pi^+(x)$.
- Use them as jump operators

$$\dot{\rho} = -i[H, \rho] + \gamma \int d^{3}x \left[2\pi^{-}\rho\pi^{+} - \{\pi^{+}\pi^{-}, \rho\} \right]$$

In momentum space,

$$\dot{\rho} = \int \frac{d^{3}\rho}{(2\pi)^{3}} \omega_{\rho} \Big(\gamma a_{\rho} \rho a_{\rho}^{\dagger} - \frac{\gamma}{2} \{ a_{\rho}^{\dagger} a_{\rho}, \rho \} - i[a_{\rho}^{\dagger} a_{\rho}, \rho] \Big)$$

By virtue of U_Λ√ω_pa_pU[†]_Λ = √ω_{Λp}a_{Λp}, the model is Lorentz covariant.

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Relativistic GKLS master equation?

arxiv:2206.05701

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Lajos Diósi Wigner Research Centre for Physics, H-1525 Budapest 114, P. O. Box 49, Hungary and Eötvös Loránd University, H-1117 Budapest, Pázmány Péter stny. 1/A

Abstract

... ... A closer look uncovers a smartly hidden defect which leaves us without Lorentz invariant Markovian master equations. They, in view of the present author, should not exist.

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Summary

Summary

• Quantum-gravity 1950's- departure from unitarity

- Standard Quantum Theory
- Relativistic approach
- Quantum measurement, collapse: not discussed
- Today: struggle to understand non-unitary dynamics learning results of non-relativistic Cat Killers

• Quantum Mechanics 1960's - departure from unitarity

- Modified Quantum Theory, to kill Cats
- Non-relativistic context
- Intrinsic link between G and quantum measurement, collapse
- Today: struggle to understand relativistic dynamics learning schemes of mainstream quantum-gravity

'... the bottle-neck of quantum-gravity is the quantum measurement postulate instead of quantum cosmology' (DICE 2008, Erice 2015)

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